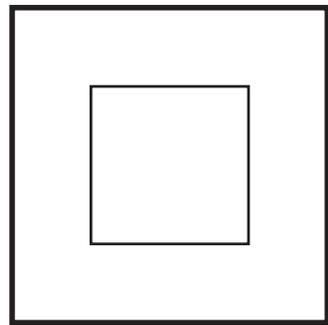


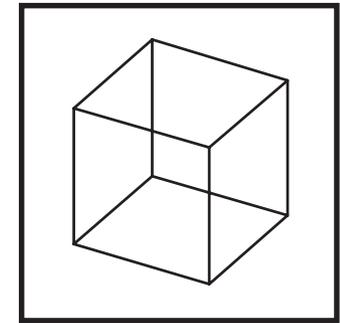
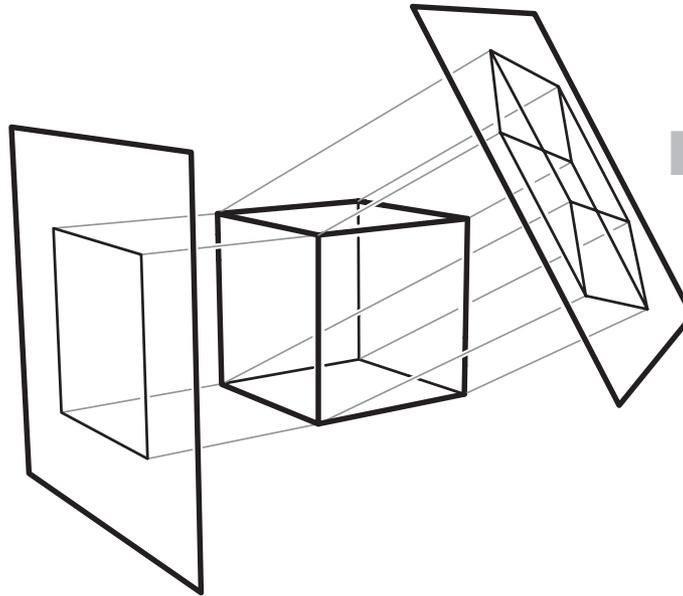
Basic Ray Tracing

CMSC 435/634

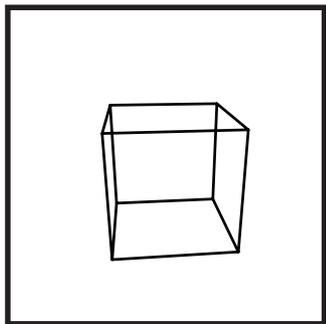
Projections



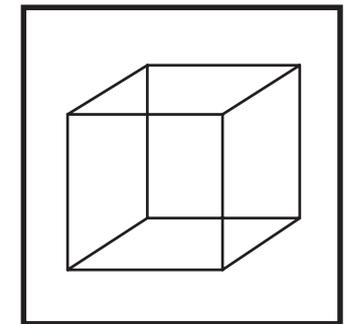
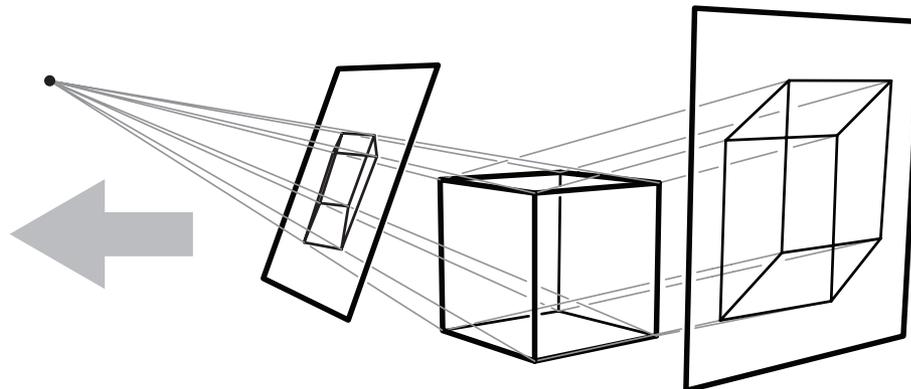
axis-aligned
orthographic



orthographic

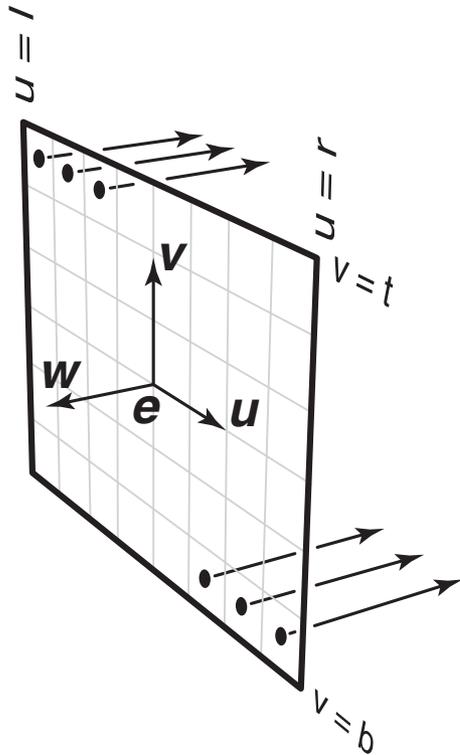


perspective

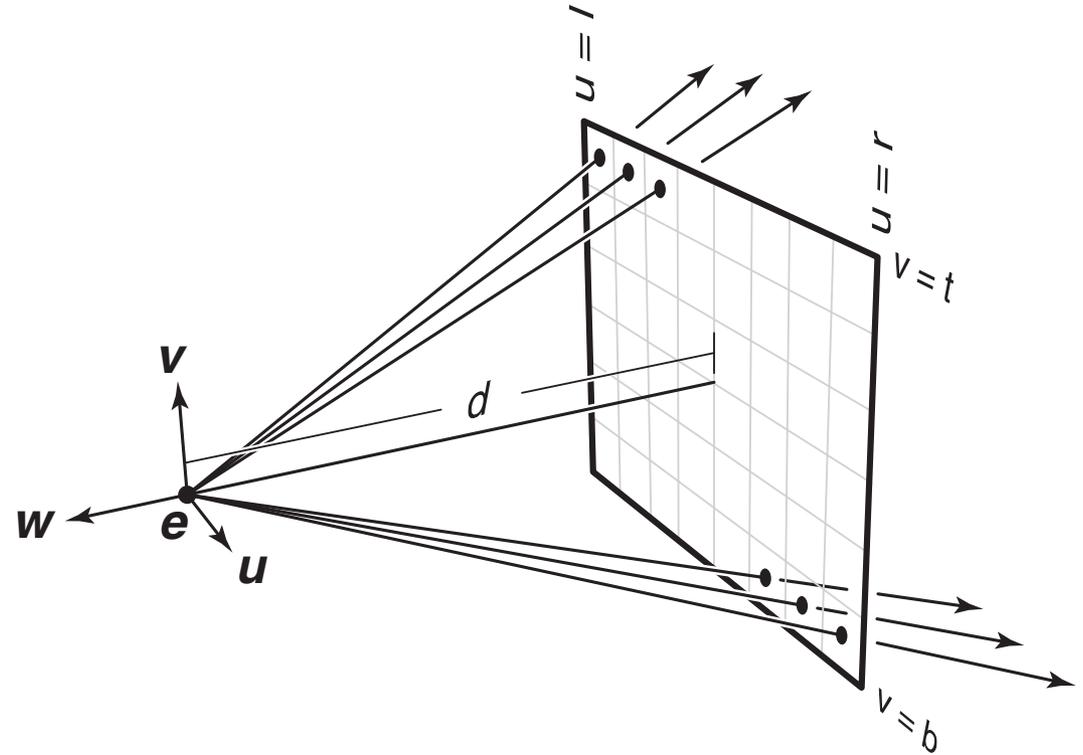


oblique

Computing Viewing Rays



Parallel projection
same direction, different origins



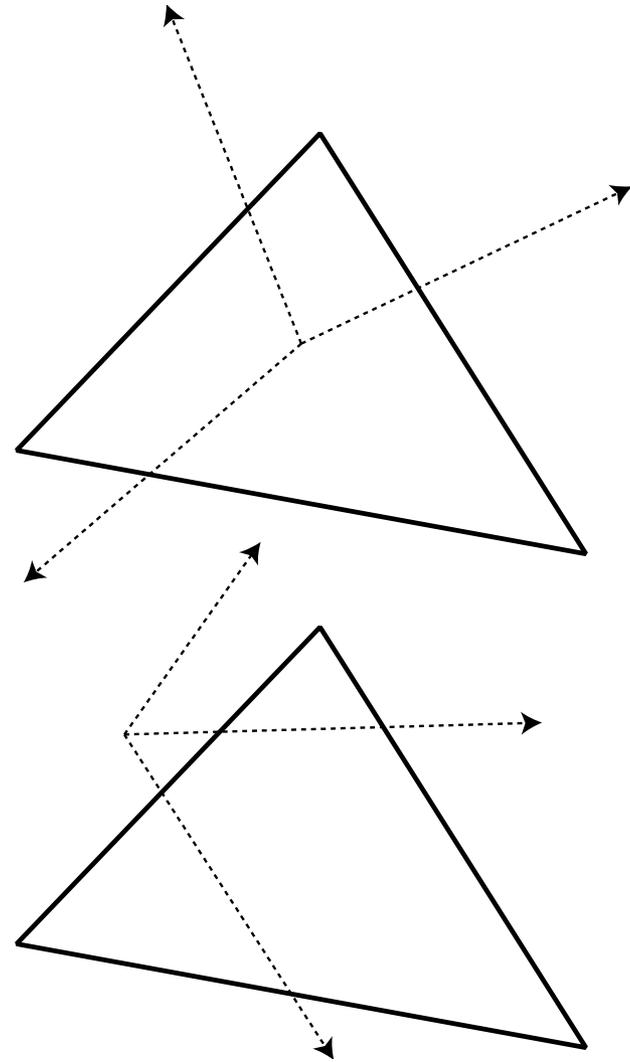
Perspective projection
same origin, different directions

Ray-Triangle Intersection

```
boolean raytri (ray r, vector p0, p1, p2, interval [t0,t1] )
{
    compute t
    if (( t < t0 ) or (t > t1))
        return ( false )
    compute  $\gamma$ 
    if (( $\gamma$  < 0 ) or ( $\gamma$  > 1))
        return ( false )
    compute  $\beta$ 
    if (( $\beta$  < 0 ) or ( $\beta + \gamma$  > 1))
        return ( false )
    return true
}
```

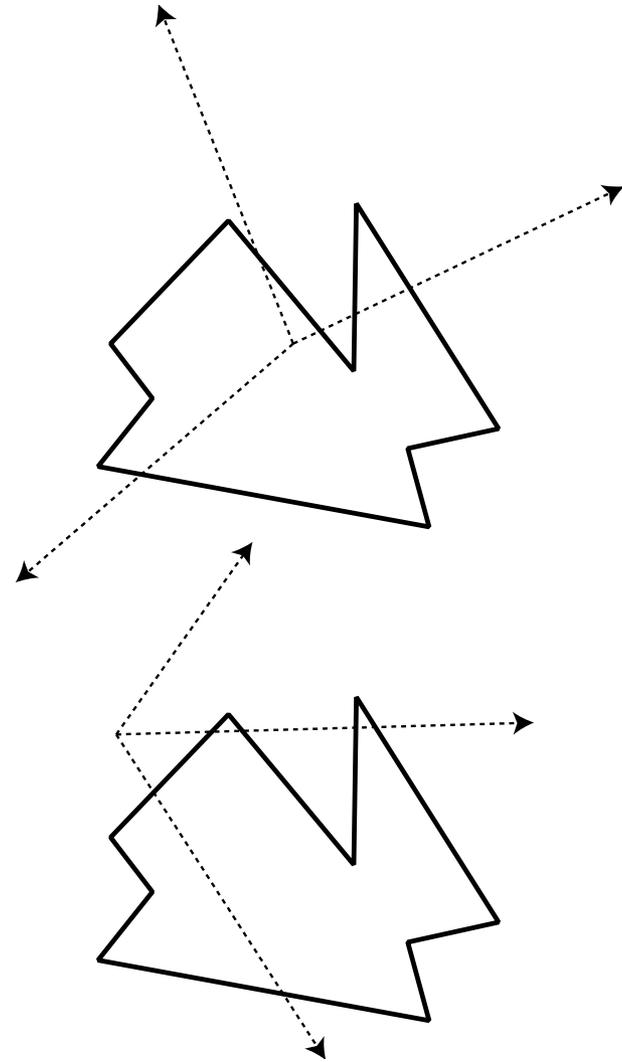
Point in Polygon?

- Is P in polygon?
- Cast ray from P to infinity
 - 1 crossing = inside
 - 0, 2 crossings = outside

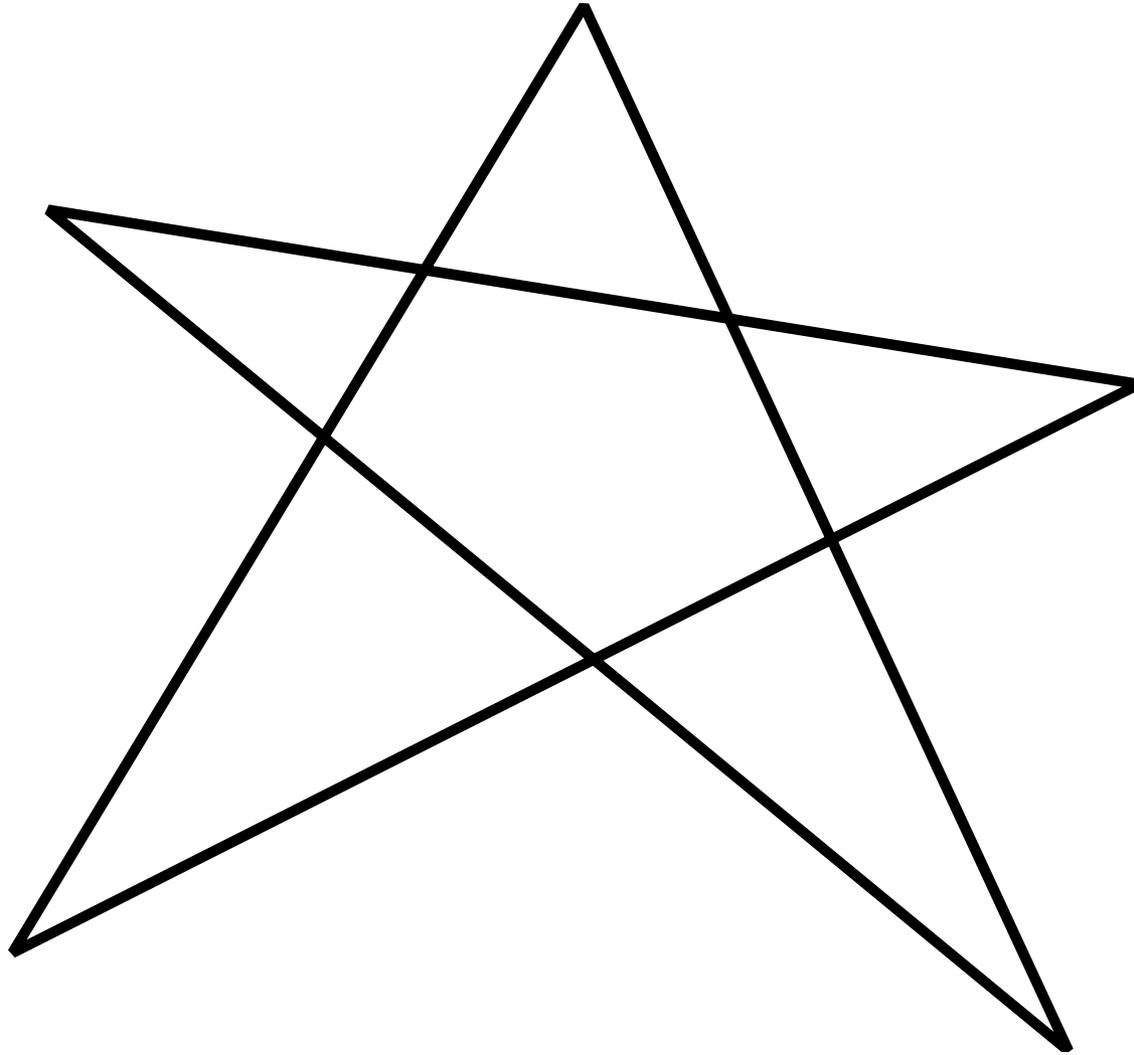


Point in Polygon?

- Is P in concave polygon?
- Cast ray from P to infinity
 - Odd crossings = inside
 - Even crossings = outside



What Happens?



Raytracing Characteristics

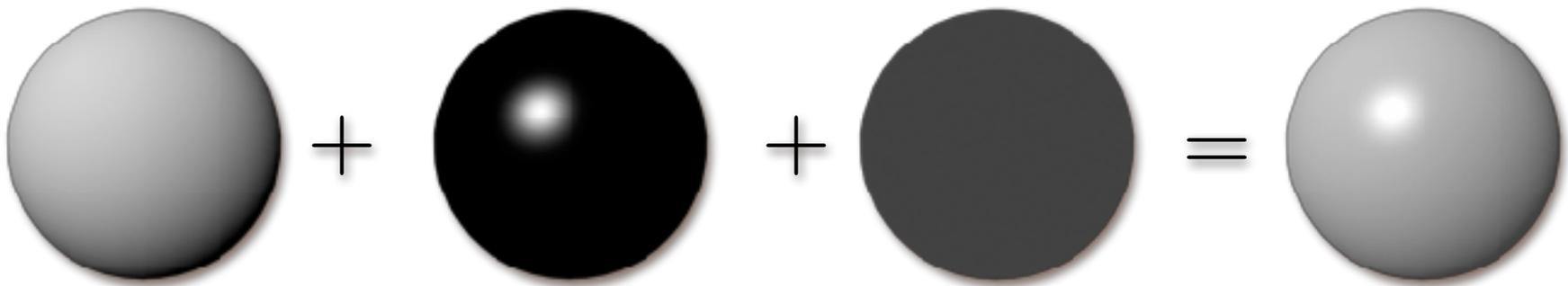
- Good
 - Simple to implement
 - Minimal memory required
 - Easy to extend
- Bad
 - Aliasing
 - Computationally intensive
 - Intersections expensive (75-90% of rendering time)
 - Lots of rays

Basic Illumination Concepts

- Terms
 - Illumination: calculating light intensity at a point (object space; equation) based loosely on physical laws
 - Shading: algorithm for calculating intensities at pixels (image space; algorithm)
- Objects
 - Light sources: light-emitting
 - Other objects: light-reflecting
- Light sources
 - Point (special case: at infinity)
 - Area

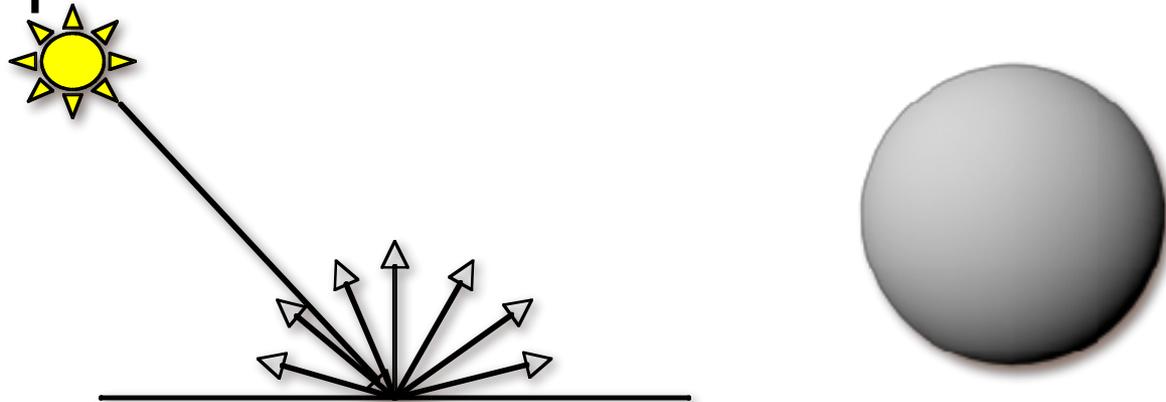
A Simple Model

- Approximate BRDF as sum of
 - A diffuse component
 - A specular component
 - A “ambient” term



Diffuse Component

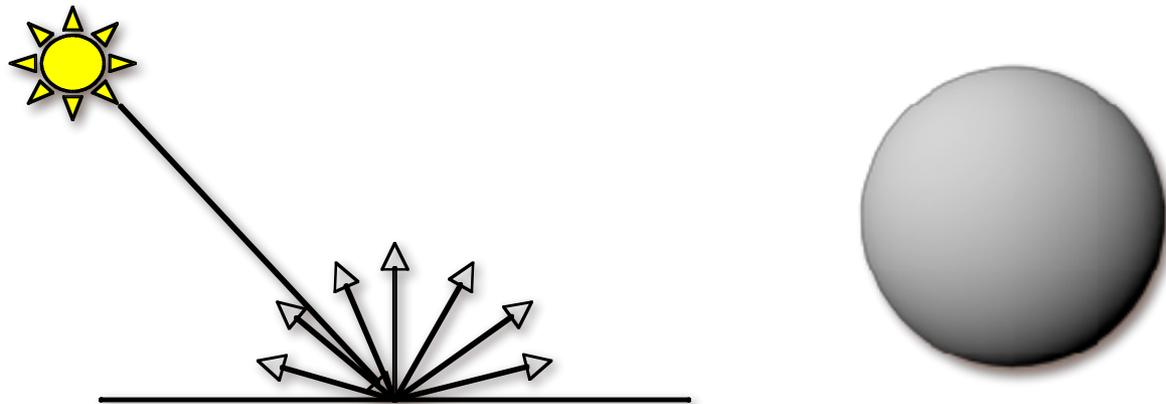
- Lambert's Law
 - Intensity of reflected light proportional to cosine of angle between surface and incoming light direction
 - Applies to “diffuse,” “Lambertian,” or “matte” surfaces
 - Independent of viewing angle
- Use as a component of non-Lambertian surfaces



Diffuse Component

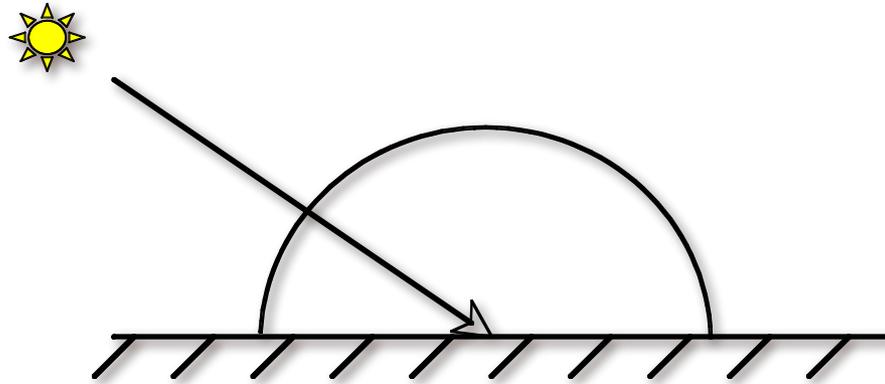
$$k_d I(\hat{\mathbf{l}} \cdot \hat{\mathbf{n}})$$

$$\max(k_d I(\hat{\mathbf{l}} \cdot \hat{\mathbf{n}}), 0)$$

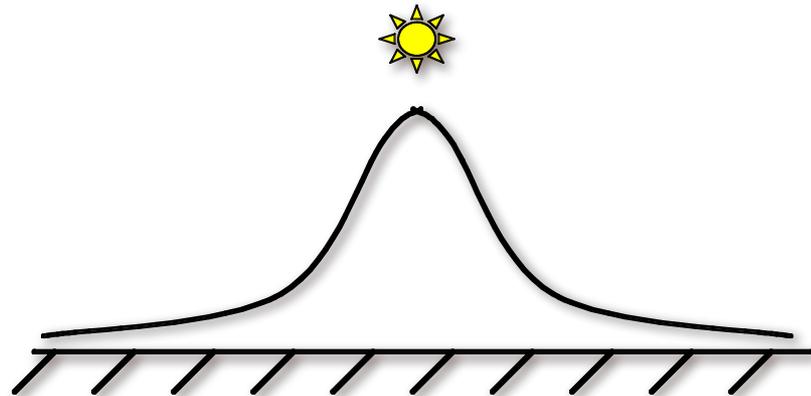


Diffuse Component

- Plot light leaving in a given direction:

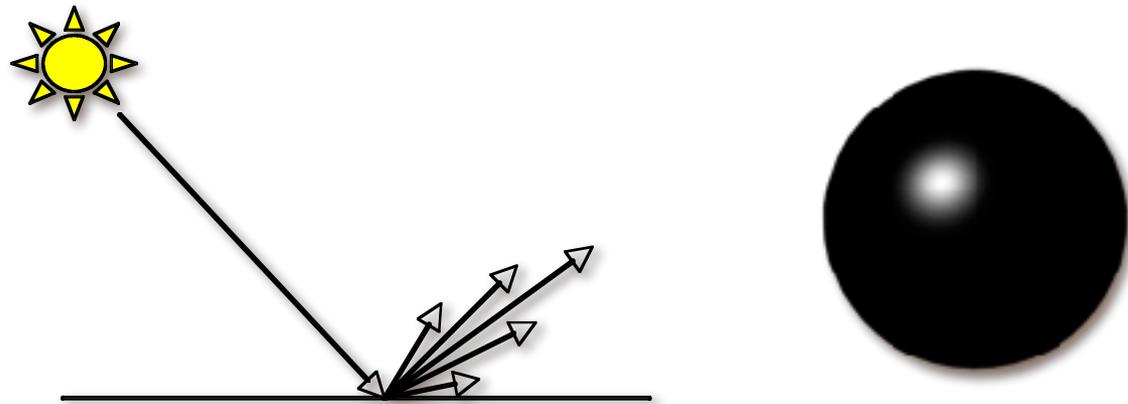


- Plot light leaving from each point on surface



Specular Component

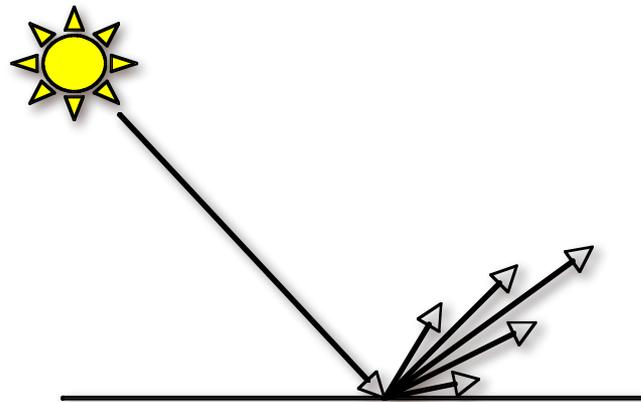
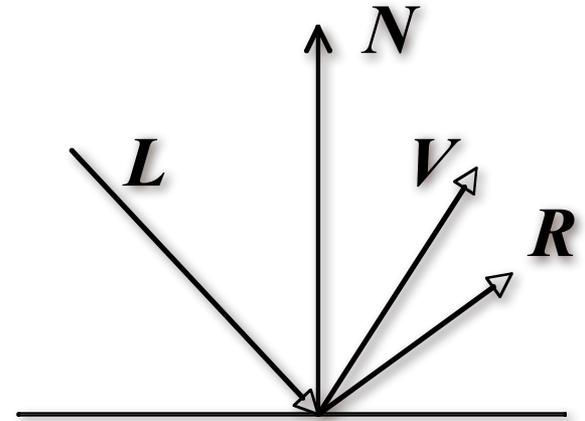
- Specular component is a mirror-like reflection
- Phong Illumination Model
 - A reasonable approximation for some surfaces
 - Fairly cheap to compute
- Depends on view direction



Specular Component

$$k_s I (\hat{\mathbf{r}} \cdot \hat{\mathbf{v}})^p$$

$$k_s I \max(\hat{\mathbf{r}} \cdot \hat{\mathbf{v}}, 0)^p$$

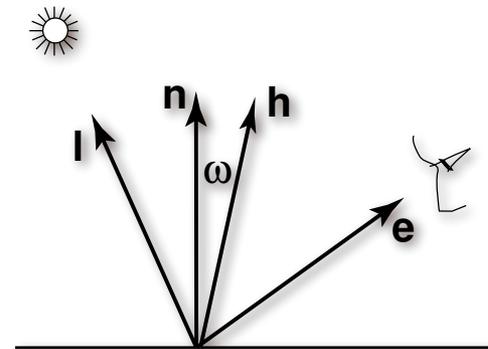
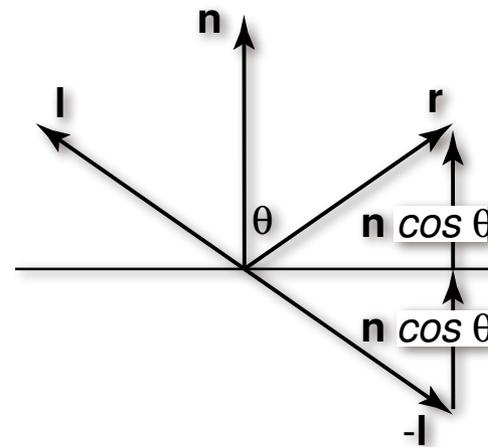


Specular Component

- Computing the reflected direction

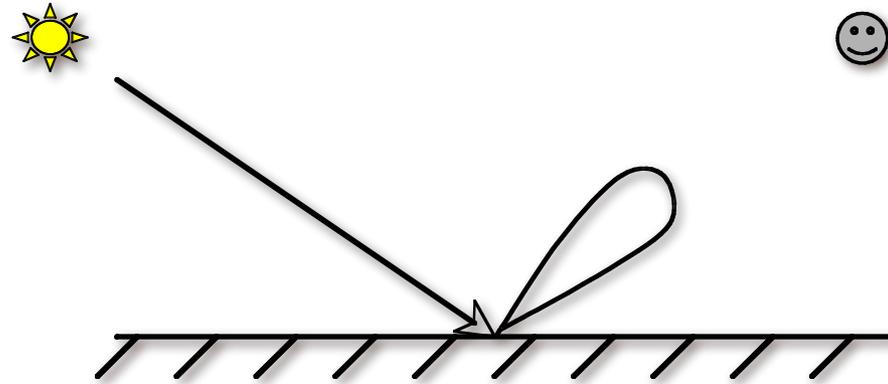
$$\hat{\mathbf{r}} = -\hat{\mathbf{l}} + 2(\hat{\mathbf{l}} \cdot \hat{\mathbf{n}})\hat{\mathbf{n}}$$

$$\hat{\mathbf{h}} = \frac{\hat{\mathbf{l}} + \hat{\mathbf{v}}}{\|\hat{\mathbf{l}} + \hat{\mathbf{v}}\|}$$

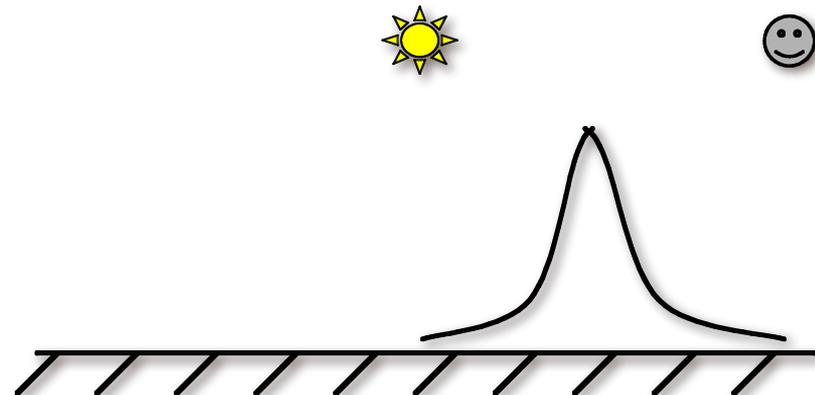


Specular Component

- Plot light leaving in a given direction:

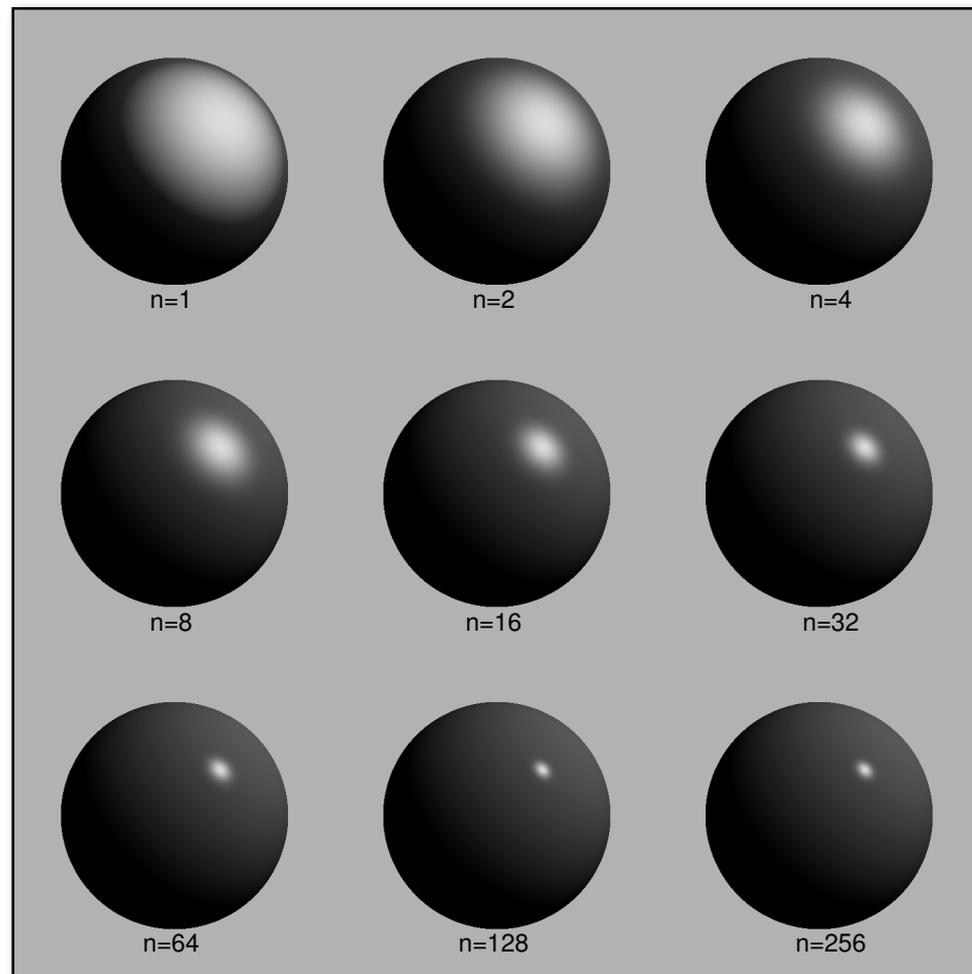


- Plot light leaving from each point on surface



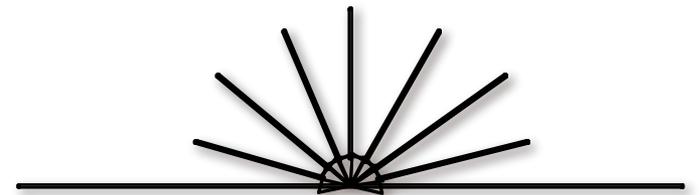
Specular Component

- Specular exponent sometimes called “roughness”



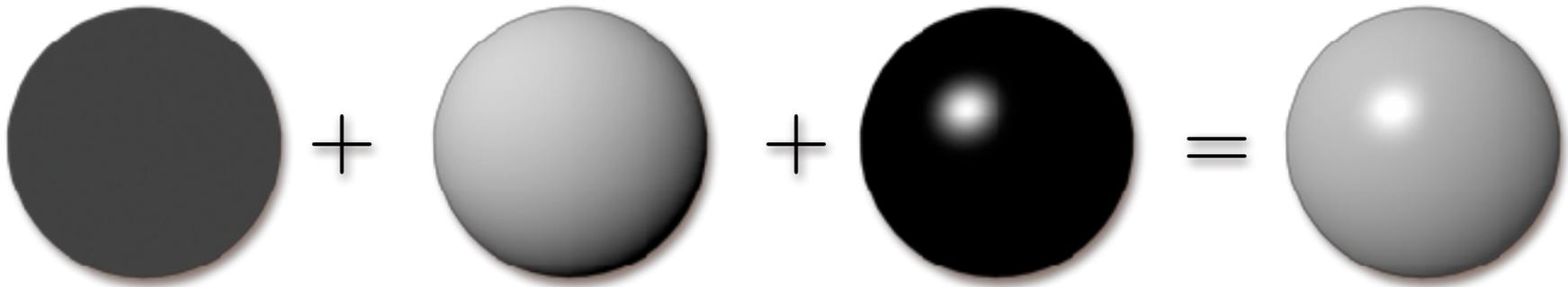
Ambient Term

- Really, its a cheap hack
- Accounts for “ambient, omnidirectional light”
- Without it everything looks like it’s in space



Summing the Parts

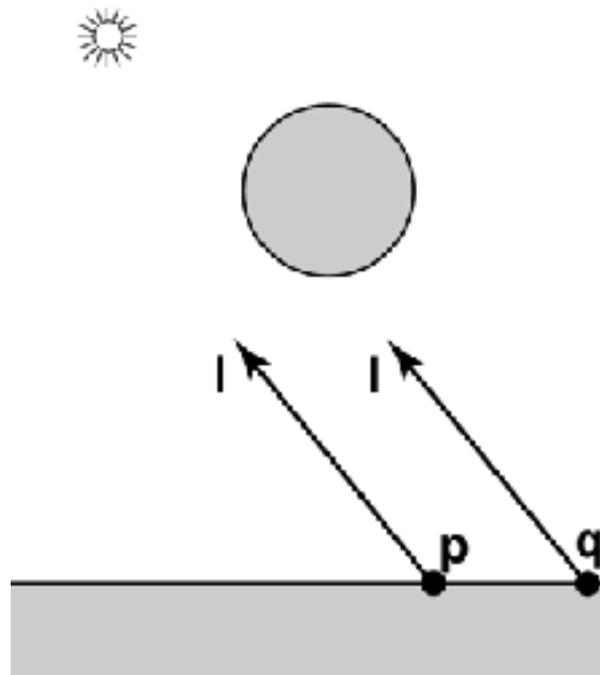
$$R = k_a I + k_d I \max(\hat{\mathbf{i}} \cdot \hat{\mathbf{n}}, 0) + k_s I \max(\hat{\mathbf{r}} \cdot \hat{\mathbf{v}}, 0)^p$$



- Recall that the k_i are by wavelength
 - RGB in practice
- Sum over all lights

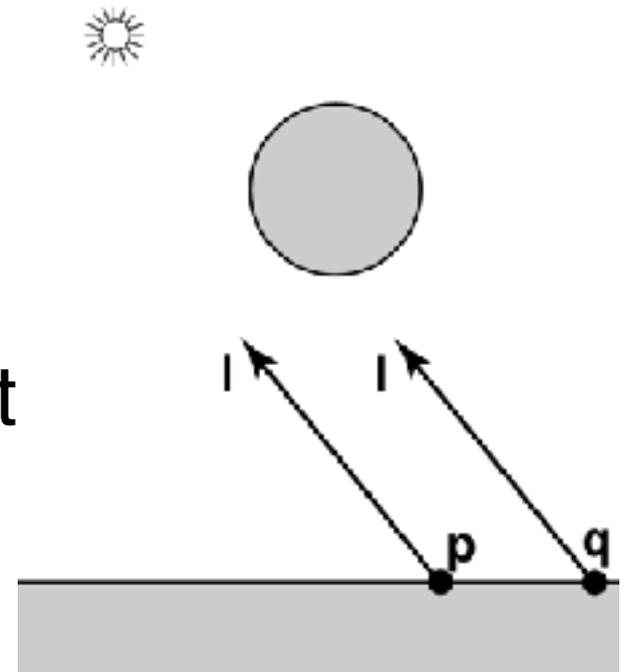
Shadows

- What if there is an object between the surface and light?

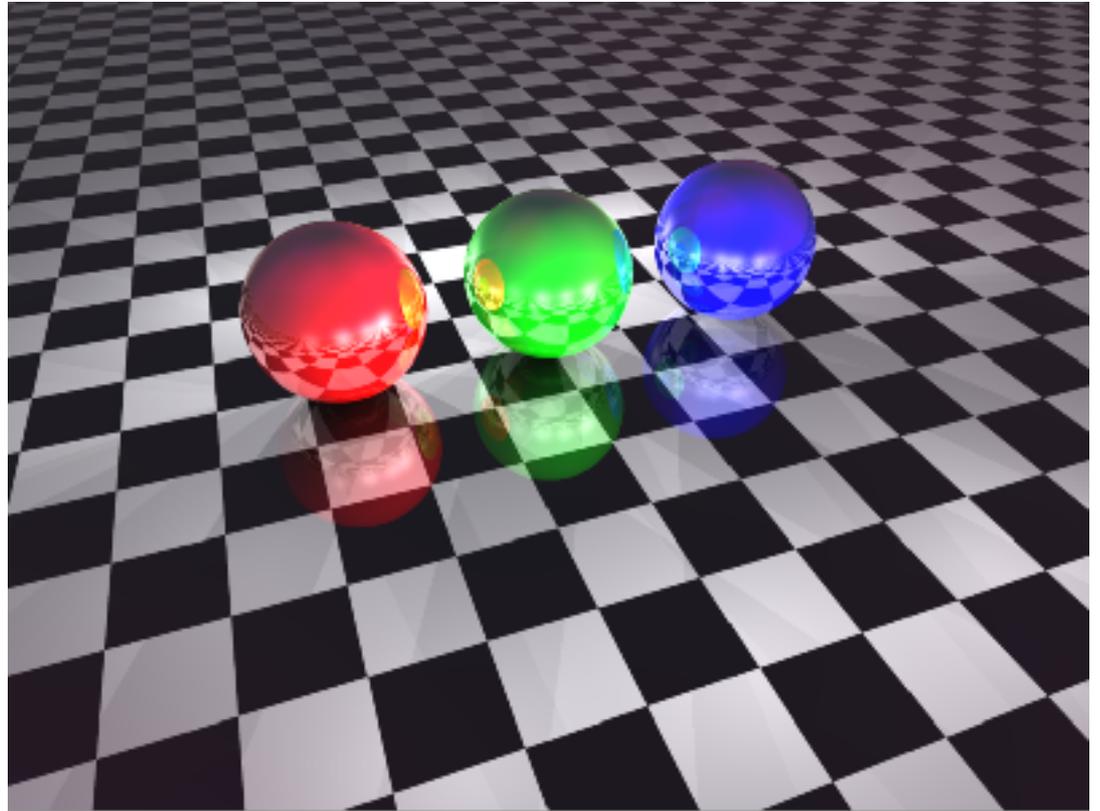


Ray Traced Shadows

- Trace a ray
 - Start = point on surface
 - End = light source
 - $t=0$ at Surface, $t=1$ at Light
 - “Bias” to avoid *surface acne*
- Test
 - $\text{Bias} \leq t \leq 1 = \text{shadow}$
 - $t < \text{Bias}$ or $t > 1 = \text{use this light}$



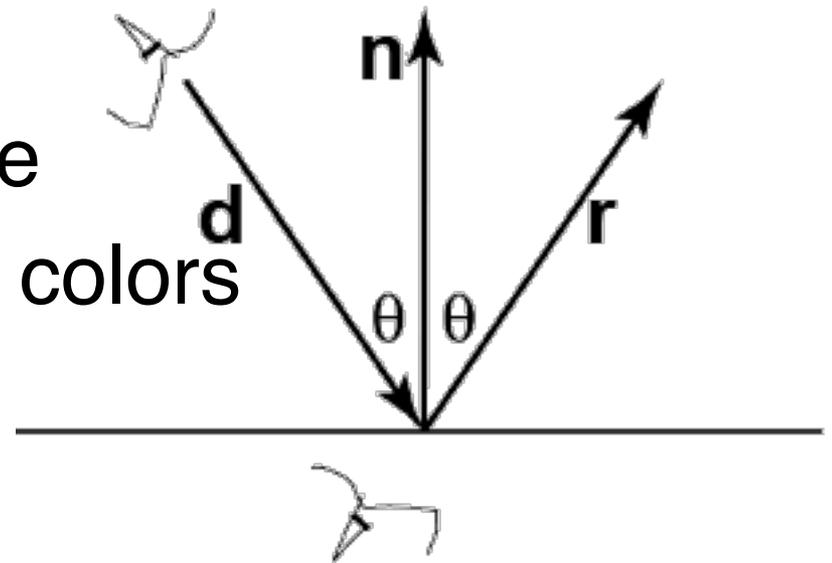
Mirror Reflection



The Dark Side of the Trees - Gilles Tran,
Spheres - Martin K. B.

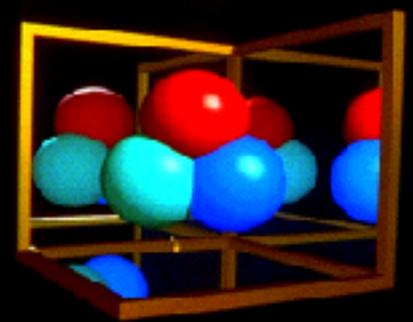
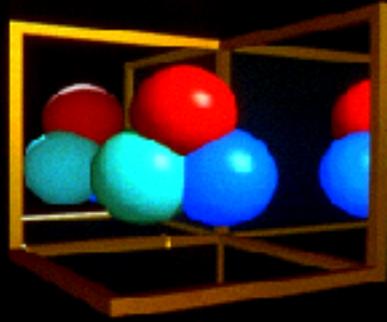
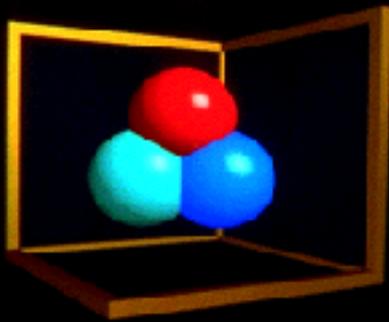
Ray Tracing Reflection

- Viewer looking in direction d sees whatever the viewer “below” the surface sees looking in direction r
- In the real world
 - Energy loss on the bounce
 - Loss different for different colors
- New ray
 - Start on surface, in reflection direction

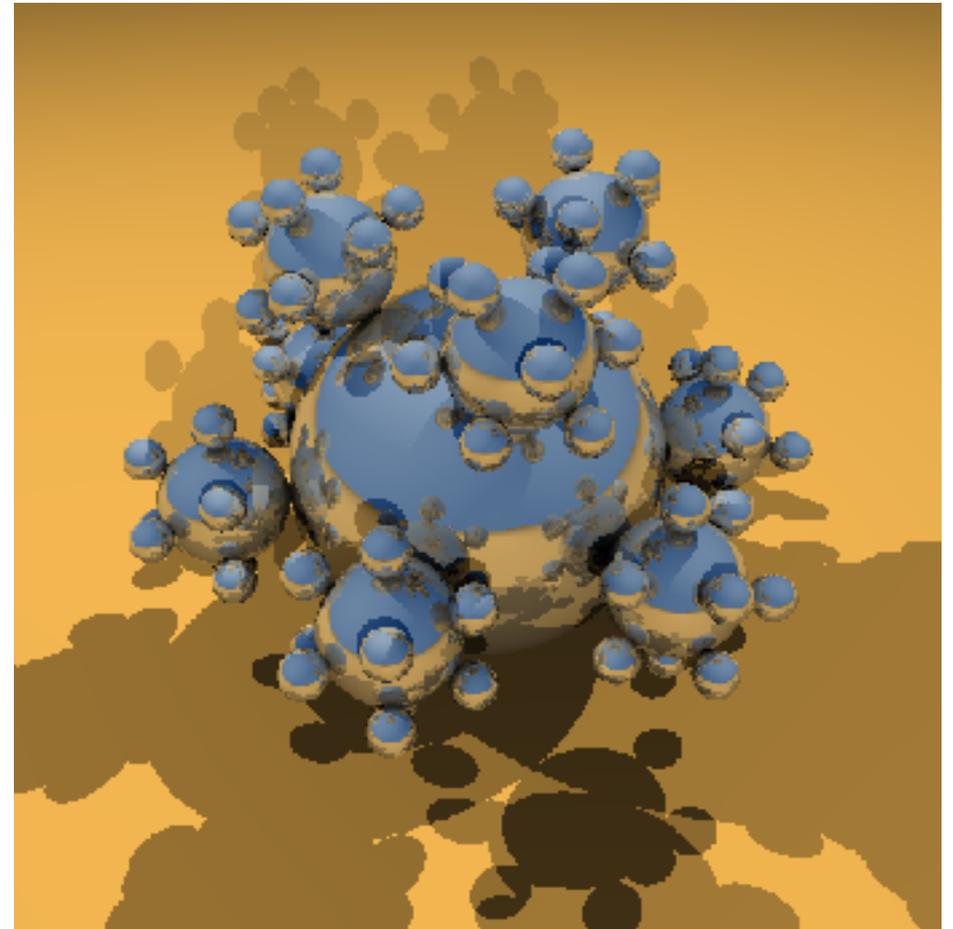


Ray Traced Reflection

- Avoid looping forever
 - Stop after n bounces
 - Stop when contribution to pixel gets too small

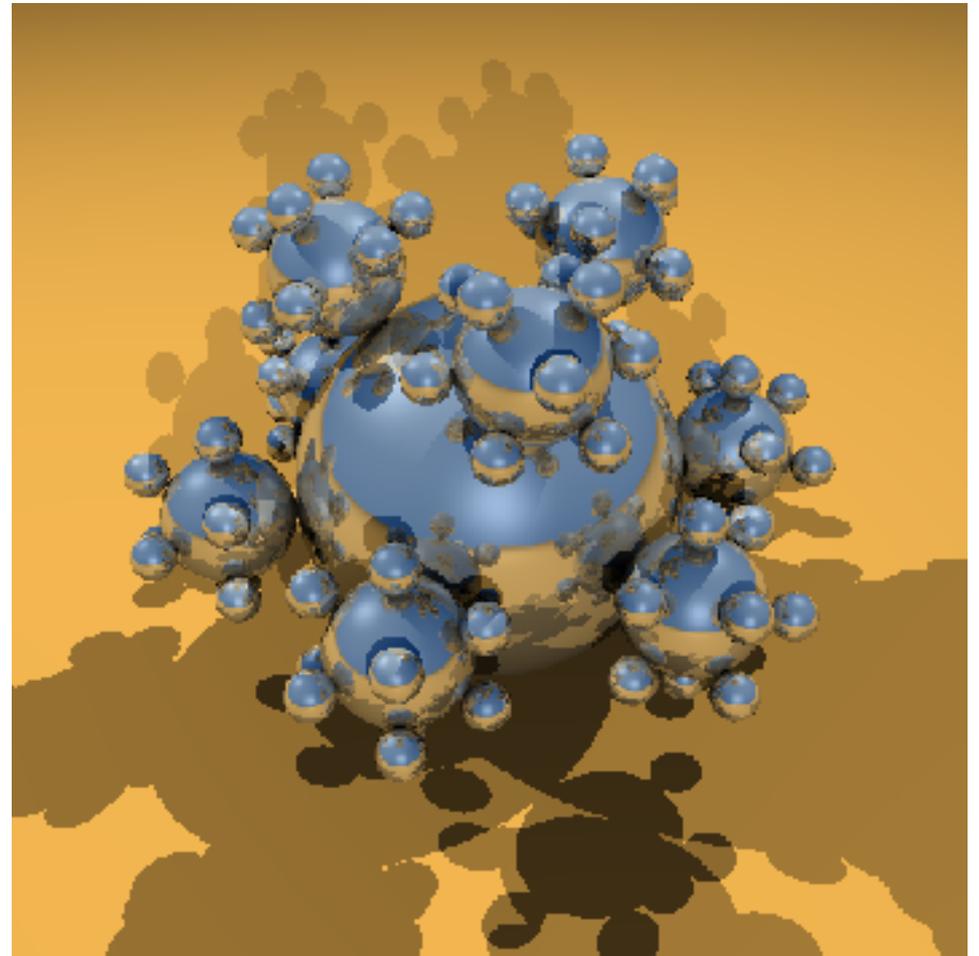


Specular vs. Mirror Reflection



Combined Specular & Mirror

- Many surfaces have both



Refraction

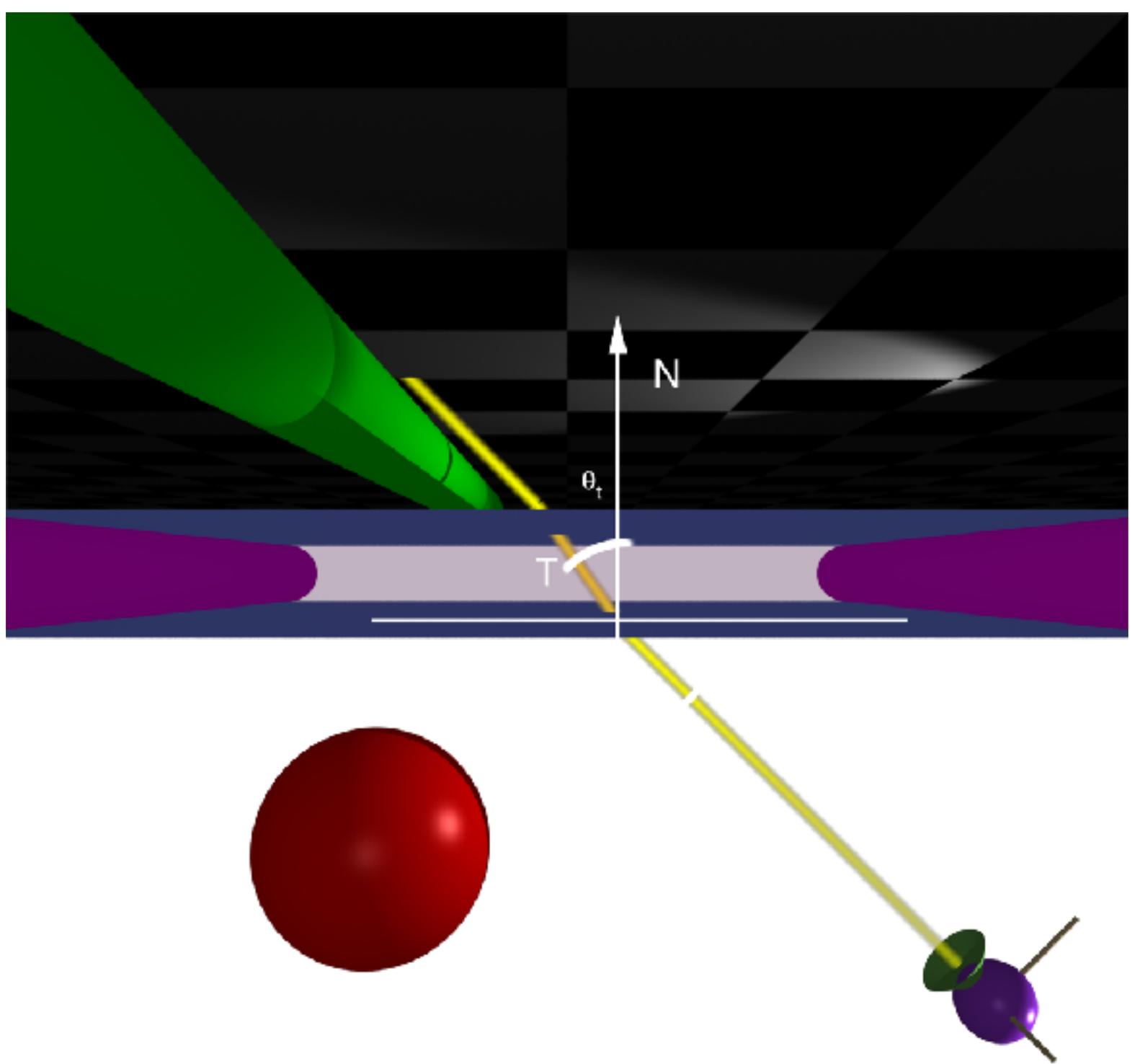




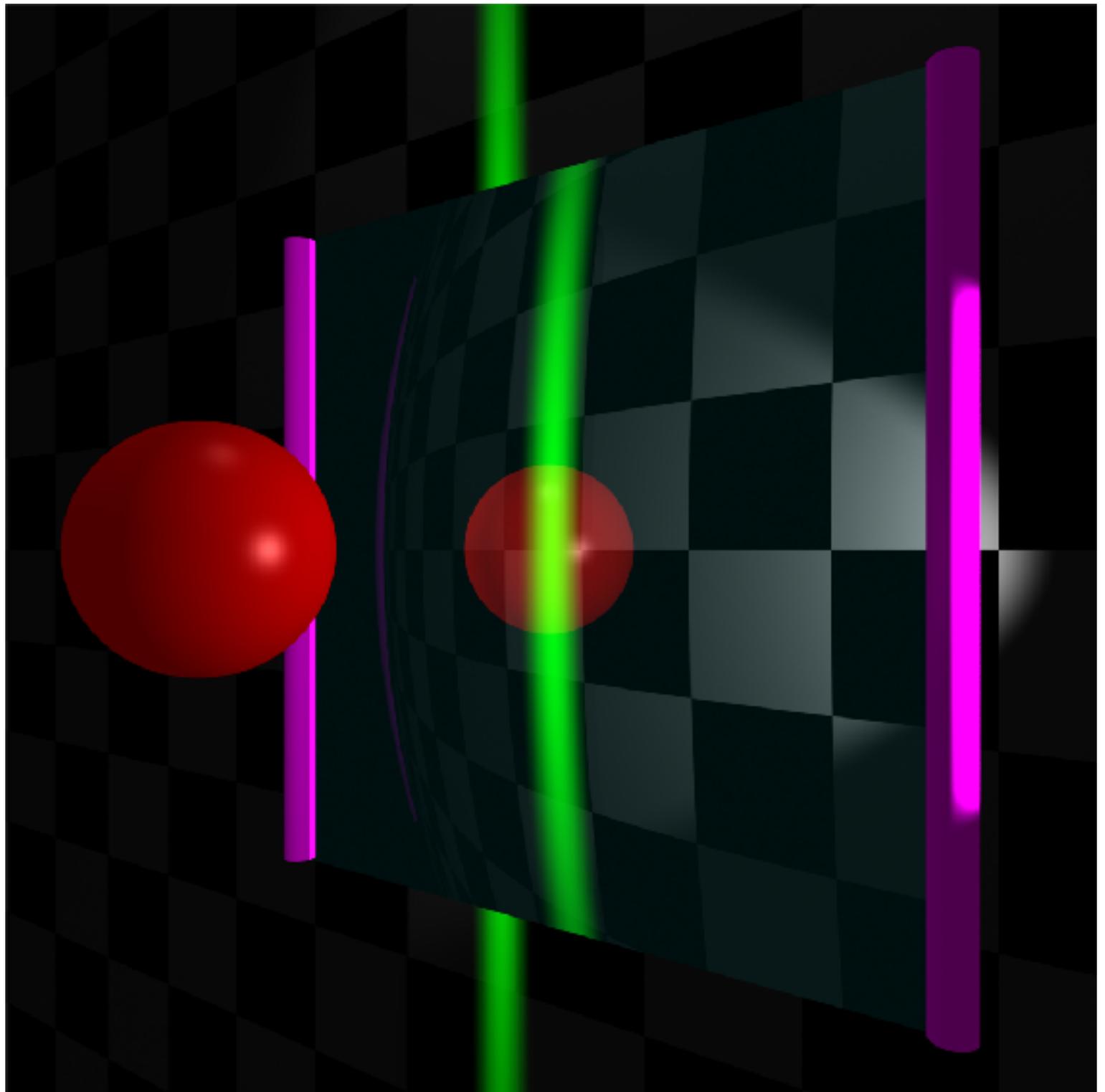
Refraction

No Refraction

Top



Front



Calculating Refraction Vector

- Snell's Law

$$n_v \sin \theta_v = n_t \sin \theta_t$$

- In terms of θ_t

$$\hat{t} = \hat{m} \sin \theta_t - \hat{n} \cos \theta_t$$

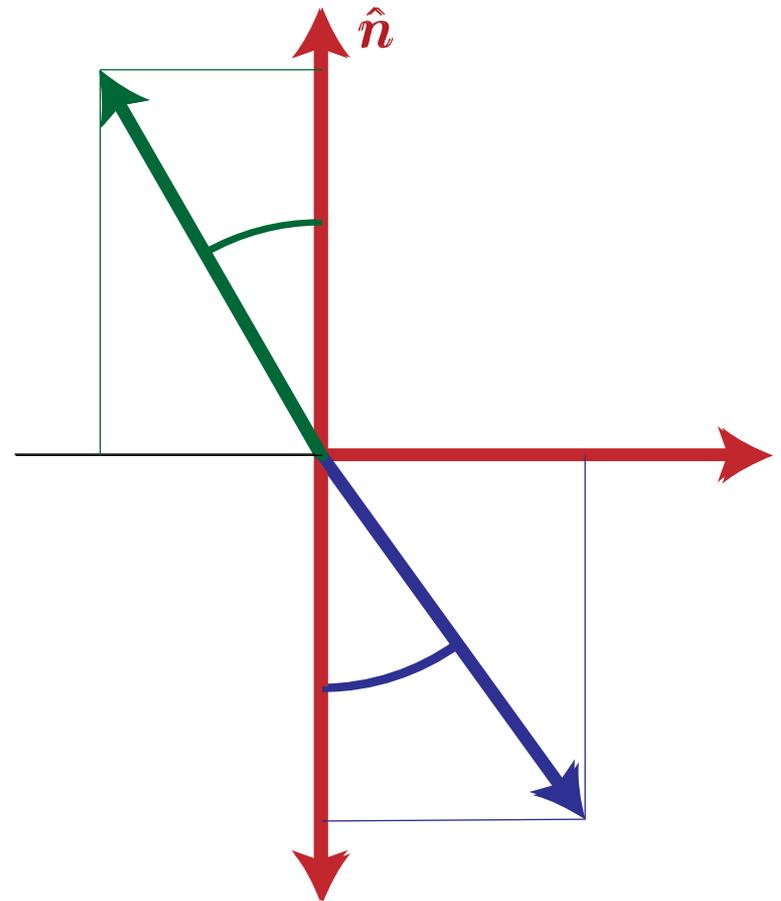
- \hat{m} term

$$\hat{m} = (\hat{n}(\hat{n} \cdot \hat{v}) - \hat{v}) / \sin \theta_v$$

$$\hat{m} \sin \theta_t$$

$$= (\hat{n}(\hat{n} \cdot \hat{v}) - \hat{v}) \sin \theta_t / \sin$$

$$= (\hat{n}(\hat{n} \cdot \hat{v}) - \hat{v}) n_v / n_t$$



Calculating Refraction Vector

- Snell's Law

$$n_v \sin \theta_v = n_t \sin \theta_t$$

- In terms of θ_t

$$\hat{t} = \hat{m} \sin \theta_t - \hat{n} \cos \theta_t$$

- \hat{n} term

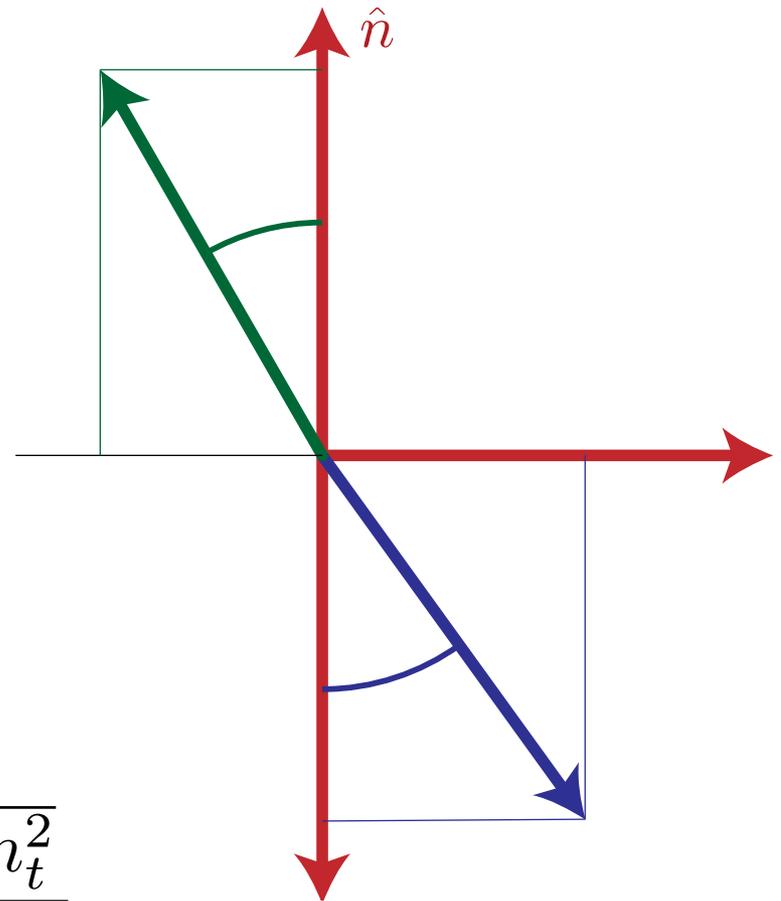
$$-\hat{n} \cos \theta_t$$

$$= -\hat{n} \sqrt{1 - \sin^2 \theta_t}$$

$$= -\hat{n} \sqrt{1 - \sin^2 \theta_v n_v^2 / n_t^2}$$

$$= -\hat{n} \sqrt{1 - (1 - \cos^2 \theta_v) n_v^2 / n_t^2}$$

$$= -\hat{n} \sqrt{1 - (1 - (\hat{n} \cdot \hat{v})^2) n_v^2 / n_t^2}$$



Calculating Refraction Vector

- Snell's Law

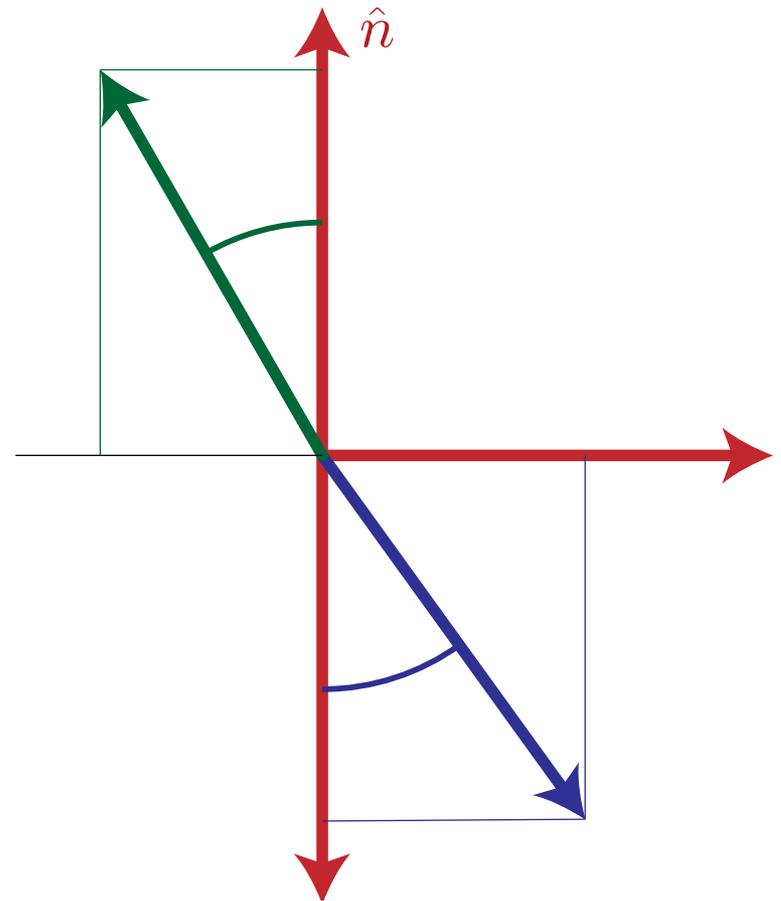
$$n_v \sin \theta_v = n_t \sin \theta_t$$

- In terms of θ_t

$$\hat{t} = \hat{m} \sin \theta_t - \hat{n} \cos \theta_t$$

- In terms of \hat{n} and \hat{v}

$$\hat{t} = (\hat{n}(\hat{n} \cdot \hat{v}) - \hat{v})n_v/n_t - \hat{n} \sqrt{1 - (1 - (\hat{n} \cdot \hat{v})^2) n_v^2/n_t^2}$$



Alpha Blending

- How much makes it through
- α = opacity
 - How much of foreground color 0-1
- $1-\alpha$ = transparency
 - How much of background color
- $\text{Foreground} * \alpha + \text{Background} * (1-\alpha)$

Refraction and Alpha

- Refraction = what direction
- α = how much
 - Often approximate as a constant
 - Better: Use Fresnel

$$F = \frac{1}{2} \left(\frac{n_v \hat{n} \cdot \hat{r} + n_t \hat{n} \cdot \hat{t}}{n_v \hat{n} \cdot \hat{r} - n_t \hat{n} \cdot \hat{t}} \right)^2 + \frac{1}{2} \left(\frac{n_v \hat{n} \cdot \hat{t} + n_t \hat{n} \cdot \hat{r}}{n_v \hat{n} \cdot \hat{t} - n_t \hat{n} \cdot \hat{r}} \right)^2$$

- Schlick approximation

$$F_0 = (n_v - n_t)^2 / (n_v + n_t)^2$$
$$F \approx F_0 + (1 - F_0)(1 - \hat{n} \cdot \hat{v})^5$$

Full Ray-Tracing

- For each pixel
 - Compute ray direction
 - Find closest surface
 - For each light
 - Shoot shadow ray
 - If not shadowed, add direct illumination
 - Shoot ray in reflection direction
 - Shoot ray in refraction direction

Dielectric

```
if (p is on a dielectric) then
  r = reflect (d, n)
  if (d.n < 0) then
    refract (d, n, n, t)
    c = -d.n
  else
    kr = kg = kb = 1
  else
    kr = exp(-alphan * t)
    kg = exp(-alphag * t)
    kb = exp(-alphan * t)
    if (refract(d, -n, 1/n t) then
      c = t.n
    else
      return k * color(p+t*r)
R0 = (n-1)^2 / (n+1)^2
R = R0 + (1-R0)(1 - c)^5
return k(R color(p + t*r) + (1-R)color(p+t*t)
```

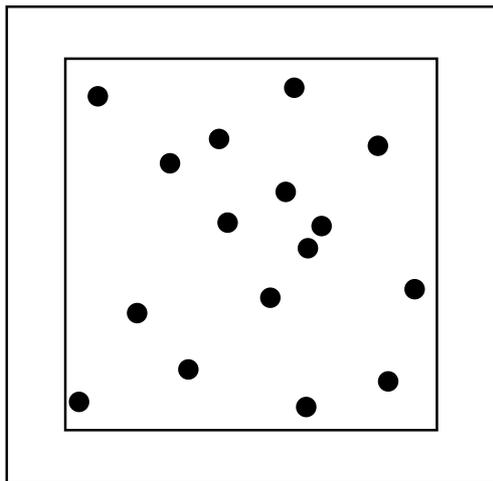
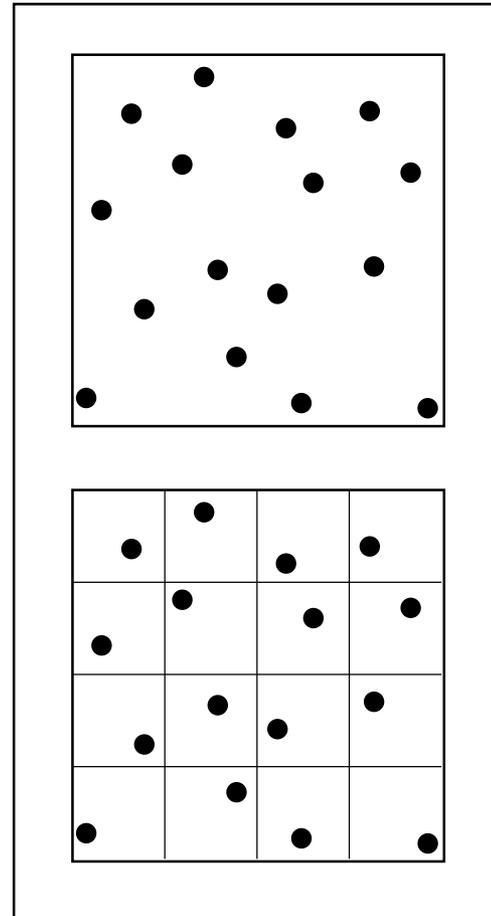
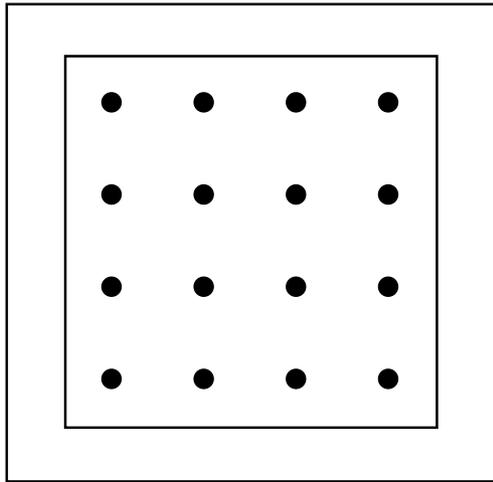
Distribution Ray Tracing

Distribution Ray Tracing

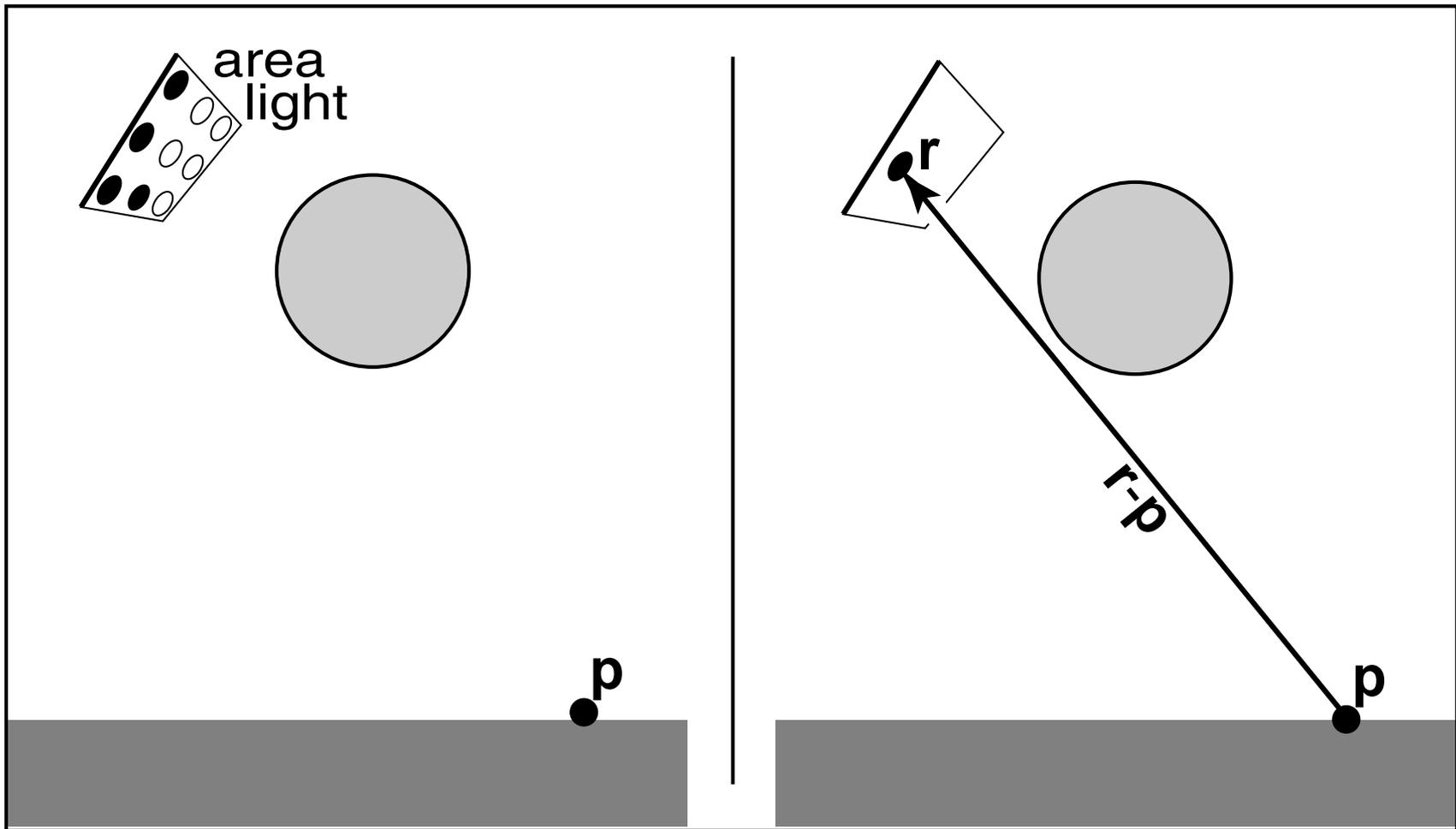
- Anti-aliasing
- Soft Shadows
- Depth of Field
- Glossy Reflection
- Motion Blur

- Turns Aliasing into Noise

Sampling



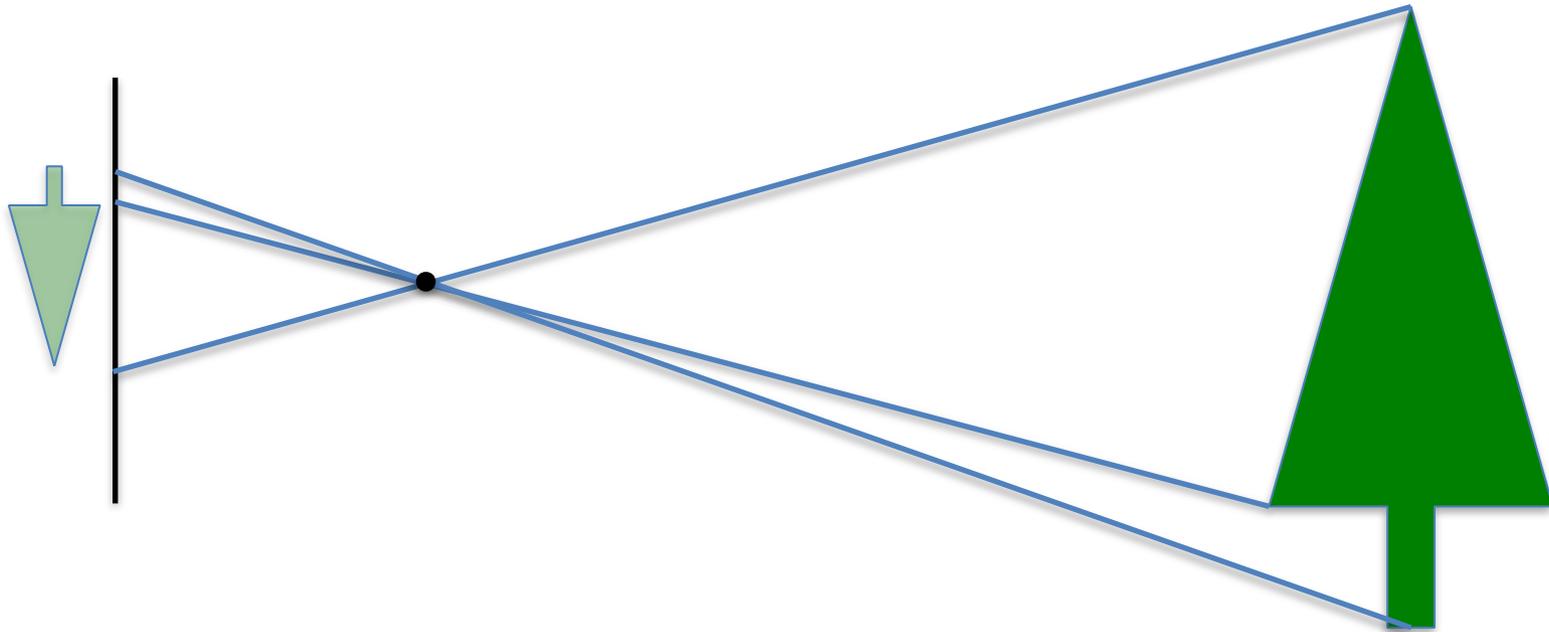
Soft Shadows



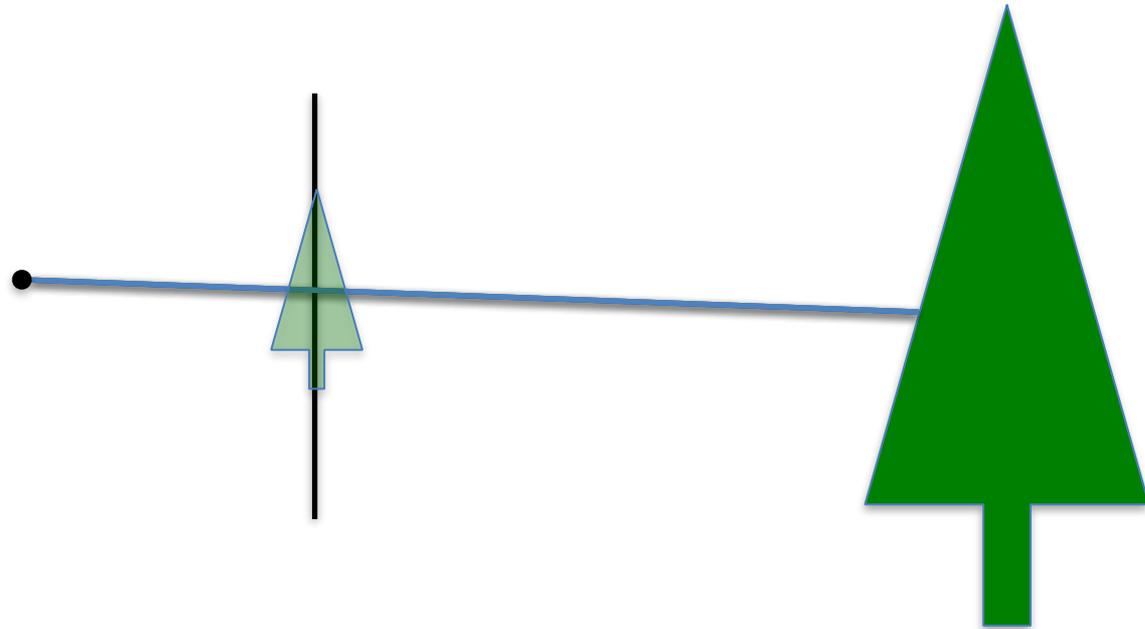
Depth of Field



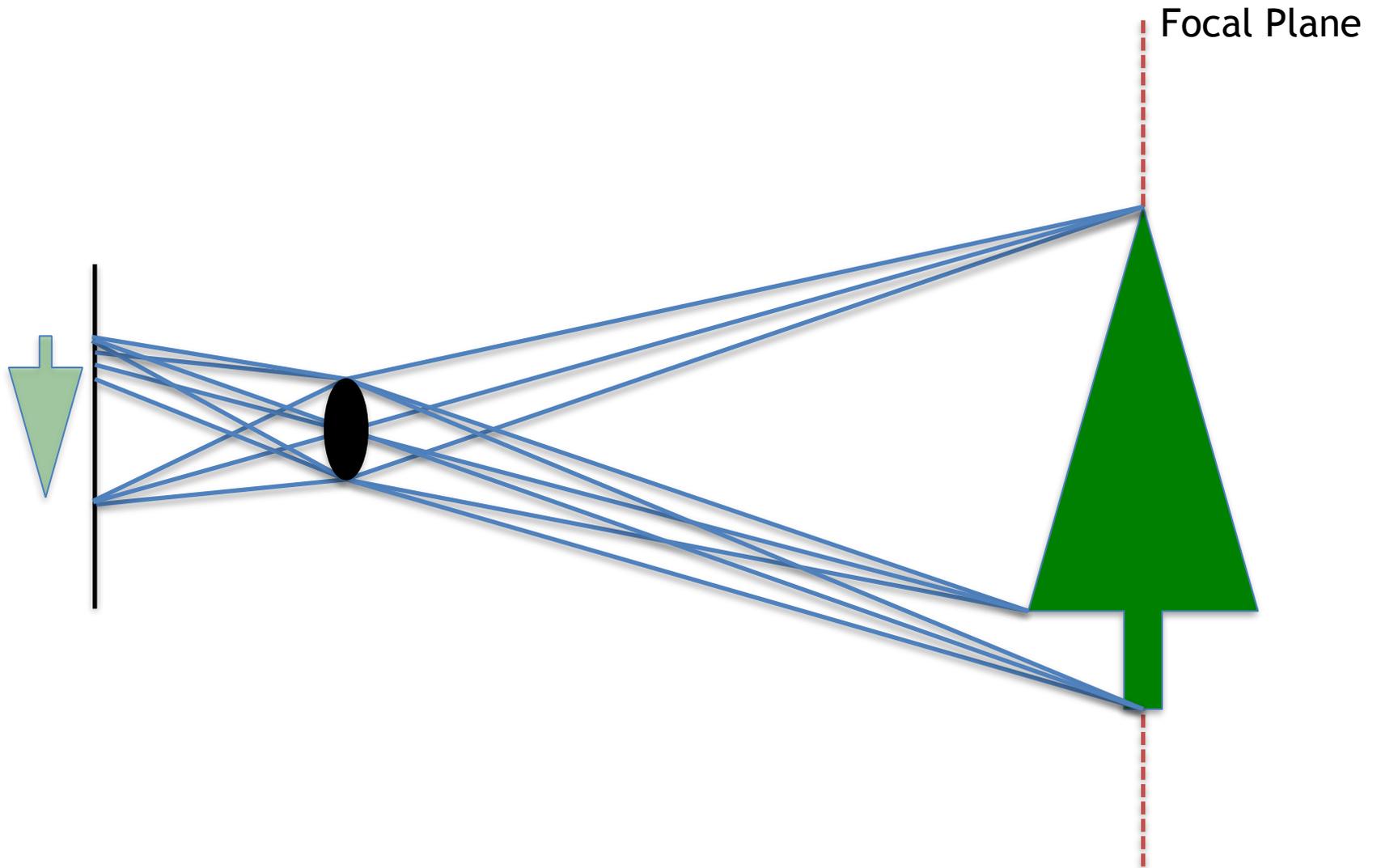
Pinhole Lens



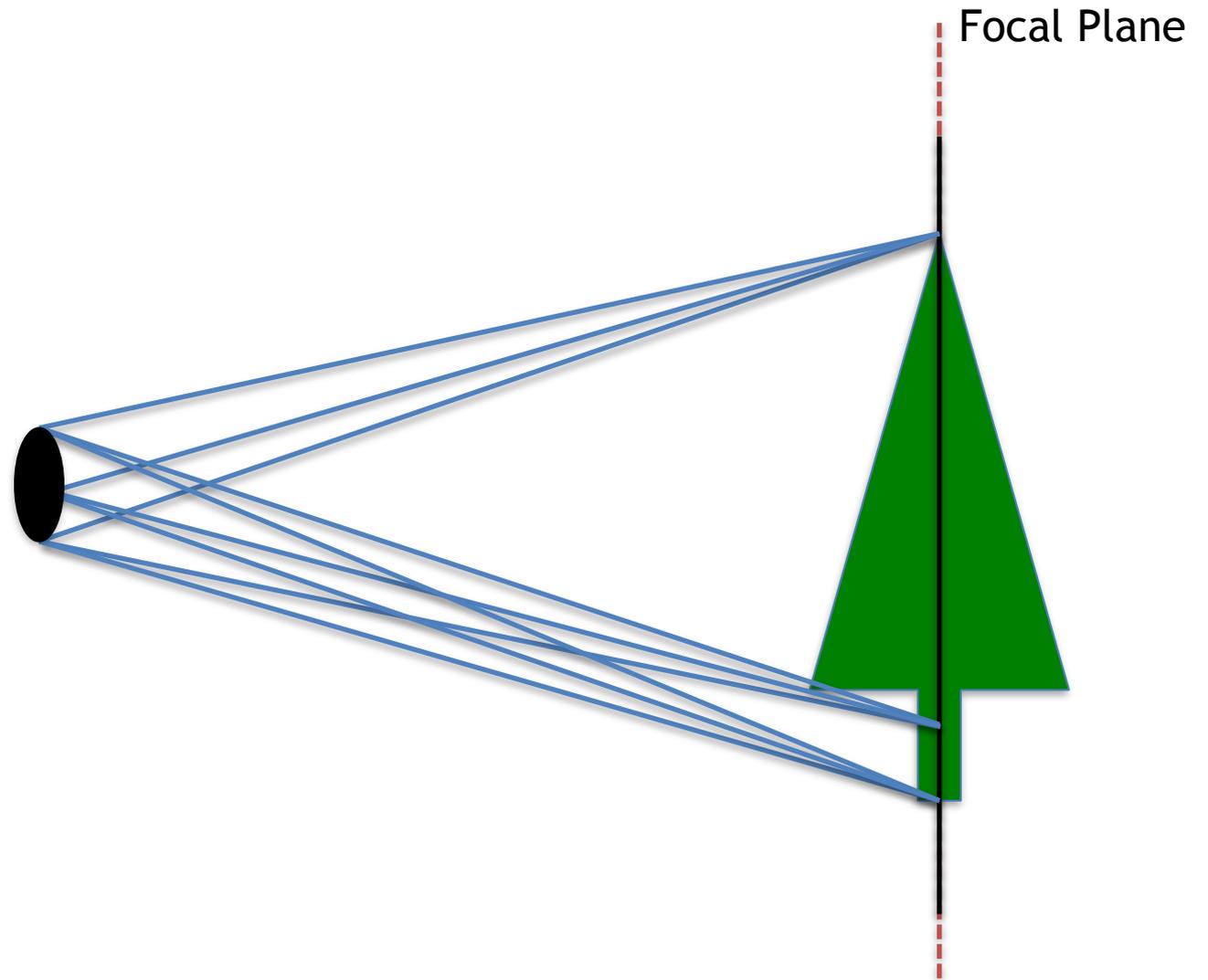
Lens Model



Real Lens



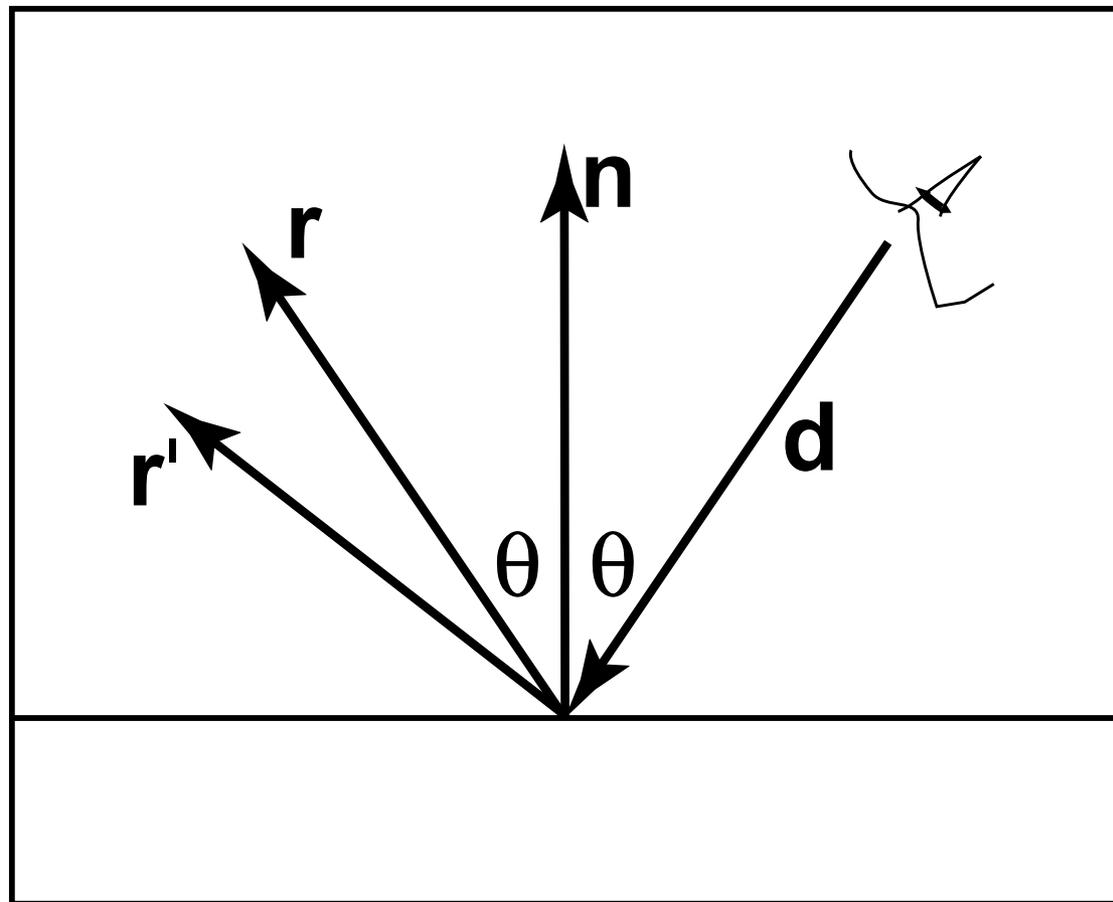
Lens Model



Ray Traced DOF

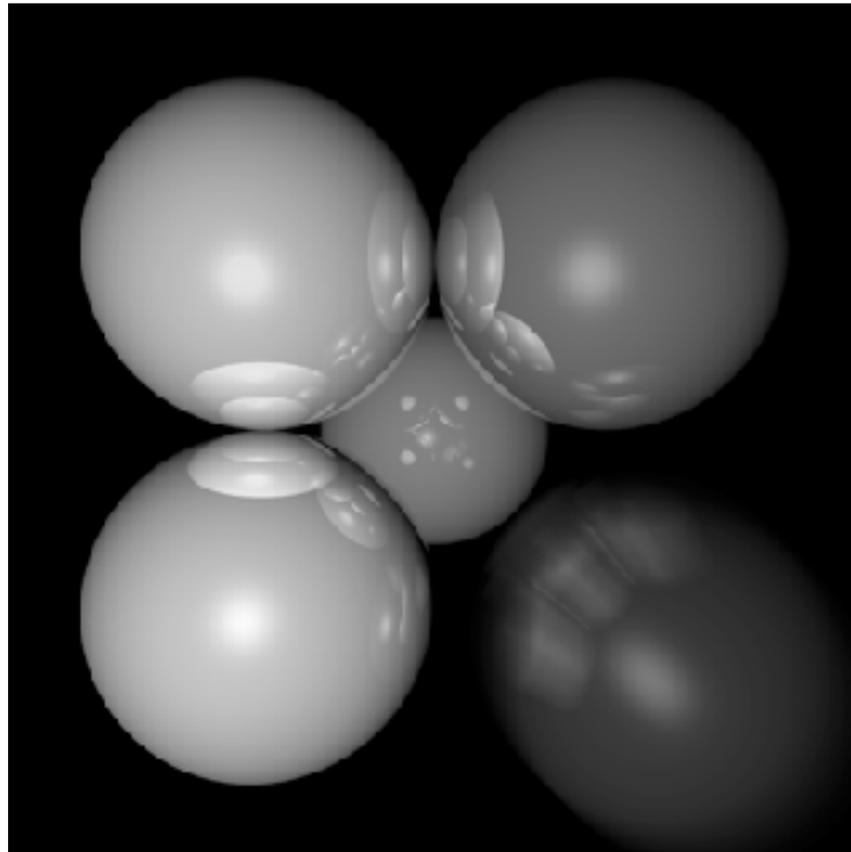
- Move image plane out to focal plane
- Jitter start position within lens aperture
 - Smaller aperture = closer to pinhole
 - Larger aperture = more DOF blur

Glossy Reflection



Motion Blur

- Things move while the shutter is open



Ray Traced Motion Blur

- Include information on object motion
- Spread multiple rays per pixel across time